## R4Photobiology A Suite of R Packages

### Pedro J. Aphalo

Department of Biosciences, University of Helsinki

and

Viikki Plant Science Center, University of Helsinki



UseR!2015, Aalborg, July 2015

#### ©2015 by Pedro J. Aphalo Department of Biosciences, University of Helsinki, Finland. http://blogs.helsinki.fi/senpep-blog/

'R4Photobiology' slide presentation by Pedro J. Aphalo is licensed under a Creative Commons Attribution-ShareAlike 4.0 International License.



## Outline

### 1 Background

- 2 Examples
- 3 Under the hood

### 4 Resources

5 Acknowledgements



# Photobiology is the study of responses to visible and ultraviolet radiation in living organisms.

- Typical tasks:
  - Describe and quantify the light environment.
  - Describe and quantify the responses of organisms to light.
  - Compute summary radiation quantities meaningful for different processes.
- Many calculations are conceptually simple but computations complex.

< ロ > < 同 > < 回 > < 回 > < 回 >



# • **Photobiology** is the study of responses to visible and ultraviolet radiation in living organisms.

- Typical tasks:
  - Describe and quantify the light environment.
  - Describe and quantify the responses of organisms to light.
  - Compute summary radiation quantities meaningful for different processes.
- Many calculations are conceptually simple but computations complex.



- Photobiology is the study of responses to visible and ultraviolet radiation in living organisms.
- Typical tasks:
  - Describe and quantify the light environment.
  - Describe and quantify the responses of organisms to light.
  - Compute summary radiation quantities meaningful for different processes.
- Many calculations are conceptually simple but computations complex.



- Photobiology is the study of responses to visible and ultraviolet radiation in living organisms.
- Typical tasks:
  - Describe and quantify the light environment.
  - Describe and quantify the responses of organisms to light.
  - Compute summary radiation quantities meaningful for different processes.
- Many calculations are conceptually simple but computations complex.



- Photobiology is the study of responses to visible and ultraviolet radiation in living organisms.
- Typical tasks:
  - Describe and quantify the light environment.
  - Describe and quantify the responses of organisms to light.
  - Compute summary radiation quantities meaningful for different processes.
- Many calculations are conceptually simple but computations complex.



- **Photobiology** is the study of responses to visible and ultraviolet radiation in living organisms.
- Typical tasks:
  - Describe and quantify the light environment.
  - Describe and quantify the responses of organisms to light.
  - Compute summary radiation quantities meaningful for different processes.
- Many calculations are conceptually simple but computations complex.

## Light-related spectra

#### What is common to all these spectra?

- A variable **wavelength**,  $\lambda$ , (or equivalent quantity) is always present
- Some other quantity either *expressed per unit wavelength* or as a ratio is also present
- Wavelength is a continuous variable, but measurements are taken at discrete and rather frequently arbitrary positions along this axis.
- The response variable is also a continuous variable and in most cases variation along the wavelength axis is also continuous.

## Light-related spectra

What is common to all these spectra?

- A variable **wavelength**,  $\lambda$ , (or equivalent quantity) is always present
- Some other quantity either *expressed per unit wavelength* or as a ratio is also present
- Wavelength is a continuous variable, but measurements are taken at discrete and rather frequently arbitrary positions along this axis.
- The response variable is also a continuous variable and in most cases variation along the wavelength axis is also continuous.

## Light-related spectra

What is common to all these spectra?

- A variable **wavelength**,  $\lambda$ , (or equivalent quantity) is always present
- Some other quantity either *expressed per unit wavelength* or as a ratio is also present
- Wavelength is a continuous variable, but measurements are taken at discrete and rather frequently arbitrary positions along this axis.
- The response variable is also a continuous variable and in most cases variation along the wavelength axis is also continuous.

# Example calculation

(Energy) irradiance

$$I_{\lambda_{\min}...\lambda_{\max}} = \int_{\lambda=\lambda_{\min}}^{\lambda=\lambda_{\max}} I(\lambda) d\lambda$$
 (1)

where  $I_{\lambda_{\min}...\lambda_{max}}$  is irradiance in  $Wm^{-2}$  for a *waveband*,  $\lambda$  is wavelength in nanometres,  $I(\lambda)$  is the spectral irradiance in  $Wm^{-2} nm^{-1}$ , and  $\lambda_{\min}$  and  $\lambda_{\max}$  are the boundaries of the waveband (region) integrated.

Although this is a simple operation, practical problems (e.g. for Excel user) are:

- **1**  $I(\lambda)$  is available only at discrete values of  $\lambda$
- **2** step size of  $\lambda$  values is variable within a single measurement
- 3  $I(\lambda)$  may need to be interpolated at the boundaries of the waveband

# Example calculation

(Energy) irradiance

$$I_{\lambda_{\min}...\lambda_{\max}} = \int_{\lambda=\lambda_{\min}}^{\lambda=\lambda_{\max}} I(\lambda) d\lambda$$
 (1)

where  $I_{\lambda_{\min}...\lambda_{max}}$  is irradiance in  $Wm^{-2}$  for a *waveband*,  $\lambda$  is wavelength in nanometres,  $I(\lambda)$  is the spectral irradiance in  $Wm^{-2} nm^{-1}$ , and  $\lambda_{\min}$  and  $\lambda_{\max}$  are the boundaries of the waveband (region) integrated.

Although this is a simple operation, practical problems (e.g. for Excel user) are:

- **1**  $I(\lambda)$  is available only at discrete values of  $\lambda$
- **2** step size of  $\lambda$  values is variable within a single measurement
- 3  $I(\lambda)$  may need to be interpolated at the boundaries of the waveband

# Example calculation

Weighted (energy) irradiance

$$I_{w} = \int_{\lambda = \lambda_{\min}}^{\lambda = \lambda_{\max}} I(\lambda) \times w(\lambda) d\lambda$$
(2)

where  $I_w$  is an *effective* irradiance in *weighted* W m<sup>-2</sup>,  $\lambda$  is wavelength in nanometres,  $I(\lambda)$  is the spectral irradiance in Wm<sup>-2</sup> nm<sup>-1</sup>,  $w(\lambda)$  is a dimensionless **spectral weighting function** (SWF).

Additional practical problems are:

- **SWFs**,  $w(\lambda)$ , can be either defined as functions or tabulated multipliers
- **2** SWFs are frequently discontinuous
- **3** SWFs outside a specified range of wavelengths are assumed to be equal to zero or to be undefined, depending on the case

# Example calculation

Weighted (energy) irradiance

$$I_{w} = \int_{\lambda = \lambda_{\min}}^{\lambda = \lambda_{\max}} I(\lambda) \times w(\lambda) d\lambda$$
<sup>(2)</sup>

where  $I_w$  is an *effective* irradiance in *weighted* Wm<sup>-2</sup>,  $\lambda$  is wavelength in nanometres,  $I(\lambda)$  is the spectral irradiance in Wm<sup>-2</sup> nm<sup>-1</sup>,  $w(\lambda)$  is a dimensionless **spectral weighting function** (SWF). Additional practical problems are:

- **SWFs**,  $w(\lambda)$ , can be either defined as functions or tabulated multipliers
- **2** SWFs are frequently discontinuous
- **3** SWFs outside a specified range of wavelengths are assumed to be equal to zero or to be undefined, depending on the case

. . .

- 「同 ト - ( 三 ト - ( 三 ト - )

# Aims: clean and friendly 'programmatic' interface

- Hide the need of interpolation from users.
- Use an integration scheme over wavelengths not affected by irregular spacing of observations.
- Use a consistent way for describing operations and operands.
- Support reproducible research approaches.
- *Reasonably* fast performance.

## Aims: help prevent user errors

- Define a class system that keeps track of different types of spectral data and their metadata.
- Provide functions for conversions among bases of expression.
- Provide functions for conversions among related quantities.
- Save data and metadata (e.g. units of expression) in same object.
- Propagate (and merge) comments and other metadata contained in objects when feasible.
- Use specialized *operators* for calculations involving spectral objects.

## Aims: make sanity checks easy

- Provide a good set of "reference" data for comparisons.
- Make it easy to combine such data to predict outcomes...
- ...e.g. light-source emission spectra and filter transmission spectra.
- Make it easy to plot different types of spectra.
- Make it easy to meaningfully annotate such plots.

### Stage 00: $\approx$ 1998 Started playing with R.

Stage 0: 2000- R became the system I use by default.

Stage 1: 2011-2012 Collected functions from existing scripts used in my research group.

Stage 2: 2012–2013 Used a "prototype" version of my photobiology package in two training events.

Stage 3: 2013 Improved performance (because we had a set of half million spectra to analyse).

Stage 4: 2014–2015 Simplified user interface for teaching and general use.

Satge 5: 2015-2016 Write book or handbook on photobiological calculations using R.

< ロ > < 同 > < 回 > < 回 > < 回 >

### Stage 00: $\approx$ 1998 Started playing with R.

### Stage 0: 2000- R became the system I use by default.

- Stage 1: 2011-2012 Collected functions from existing scripts used in my research group.
- Stage 2: 2012–2013 Used a "prototype" version of my photobiology package in two training events.
- Stage 3: 2013 Improved performance (because we had a set of half million spectra to analyse).
- Stage 4: 2014–2015 Simplified user interface for teaching and general use.
- Satge 5: 2015-2016 Write book or handbook on photobiological calculations using R.

< ロ > < 同 > < 回 > < 回 > < 回 >

Stage 00:  $\approx$  1998 Started playing with R.

Stage 0: 2000- R became the system I use by default.

# Stage 1: 2011-2012 Collected functions from existing scripts used in my research group.

Stage 2: 2012-2013 Used a "prototype" version of my photobiology package in two training events.

# Stage 3: 2013 Improved performance (because we had a set of half million spectra to analyse).

Stage 4: 2014–2015 Simplified user interface for teaching and general use.

Satge 5: 2015-2016 Write book or handbook on photobiological calculations using R.

< ロ > < 同 > < 回 > < 回 > < 回 > <

Stage 00:  $\approx$  1998 Started playing with R.

Stage 0: 2000- R became the system I use by default.

Stage 1: 2011-2012 Collected functions from existing scripts used in my research group.

# Stage 2: 2012–2013 Used a "prototype" version of my photobiology package in two training events.

- Stage 3: 2013 Improved performance (because we had a set of half million spectra to analyse).
- Stage 4: 2014-2015 Simplified user interface for teaching and general use.

Satge 5: 2015-2016 Write book or handbook on photobiological calculations using R.

Stage 00:  $\approx$  1998 Started playing with R.

Stage 0: 2000- R became the system I use by default.

- Stage 1: 2011-2012 Collected functions from existing scripts used in my research group.
- Stage 2: 2012–2013 Used a "prototype" version of my photobiology package in two training events.
- Stage 3: 2013 Improved performance (because we had a set of half million spectra to analyse).

Stage 4: 2014–2015 Simplified user interface for teaching and general use.

Satge 5: 2015-2016 Write book or handbook on photobiological calculations using R.

Stage 00:  $\approx$  1998 Started playing with R.

Stage 0: 2000- R became the system I use by default.

- Stage 1: 2011-2012 Collected functions from existing scripts used in my research group.
- Stage 2: 2012–2013 Used a "prototype" version of my photobiology package in two training events.
- Stage 3: 2013 Improved performance (because we had a set of half million spectra to analyse).
- Stage 4: 2014-2015 Simplified user interface for teaching and general use.

Satge 5: 2015-2016 Write book or handbook on photobiological calculations using R.

< ロ > < 同 > < 回 > < 回 > < 回 > <

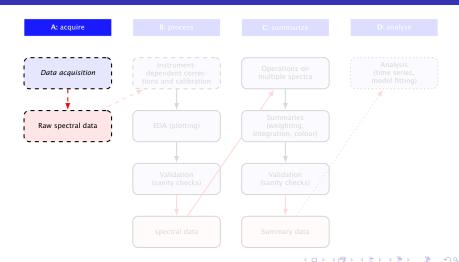
Stage 00:  $\approx$  1998 Started playing with R.

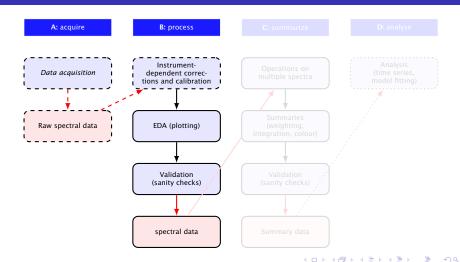
Stage 0: 2000- R became the system I use by default.

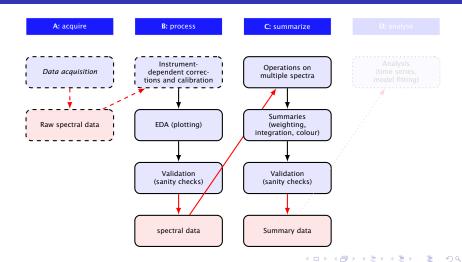
- Stage 1: 2011-2012 Collected functions from existing scripts used in my research group.
- Stage 2: 2012–2013 Used a "prototype" version of my photobiology package in two training events.
- Stage 3: 2013 Improved performance (because we had a set of half million spectra to analyse).
- Stage 4: 2014-2015 Simplified user interface for teaching and general use.

Satge 5: 2015-2016 Write book or handbook on photobiological calculations using R.

< ロ > < 同 > < 三 > < 三 > < 三 > <



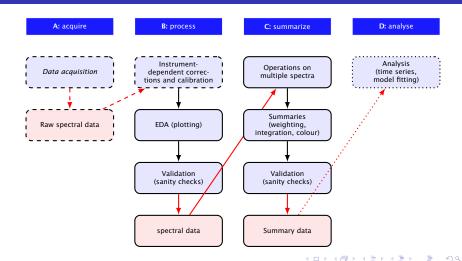




© Pedro J. Aphalo (Univ. of Helsinki)

R4Photobiology

UseR!2015, Aalborg, July 2015 12 / 33



© Pedro J. Aphalo (Univ. of Helsinki)

R4Photobiology

UseR!2015, Aalborg, July 2015 12 / 33

— Examples

# B: Example data

Spectral irradiance of sunlight

sun.spct							
##		w.length	s.e.irrad				
##	1:	293	2.609665e-06				
##	2:	294	6.142401e-06				
##	3:	295	2.176175e-05				
##							
##	506:	798	4.236281e-01				
##	507:	799	4.185850e-01				
##	508:	800	4.069055e-01				

w.length = wavelength in namometres ( $10^{-9}$  m, 1 nm = 10 Å) s.e.irrad = spectral energy irradiance in  $W m^{-2} nm^{-1}$ 

## B: Example data

Transmittance of a glass filter from Schott

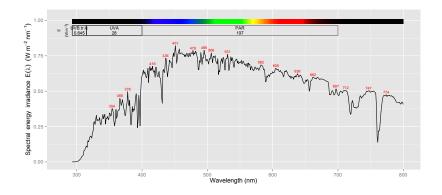
gg400.spct						
##		w.length				
##	1:	200	1e-05			
##	2:	210	1e-05			
##	3:	220	1e-05			
##						
##	178:	5050	1e-05			
##	179:	5100	1e-05			
##	180:	5150	1e-05			

w.length = wavelength in namometres ( $10^{-9}$  m, 1 nm = 10 Å) Tfr = spectral transmittance as a fraction of one

# B: plot methods for spectra

A sunlight spectrum

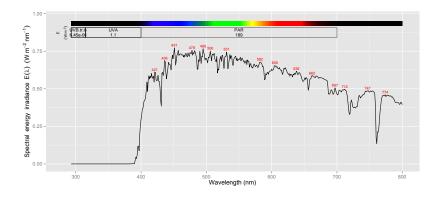
#### plot(sun.spct)



## C: Operators for spectra

e.g. using operators to simulate spectral irradiance under a filter

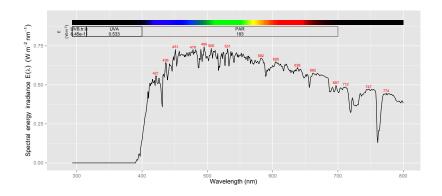
#### plot(sun.spct \* gg400.spct)



## C: Operators for spectra

e.g. using operators to simulate the effect of two stacked filters

#### plot(sun.spct \* gg400.spct^2)



© Pedro J. Aphalo (Univ. of Helsinki)

#### — Examples

## C: Summaries of spectra

summary, wavelength range, equivalent color definition

```
summary(sun.spct)
## wavelength ranges from 293 to 800 nm
  largest wavelength step size is 1 nm
##
  spectral irradiance ranges from 2.61e-06 to 0.8205 W m-2 nm-1
##
## energy irradiance is 269.1 W m-2
range(sun.spct)
## [1] 293 800
color(sun.spct)
  source CMF source CC
##
##
    "#544F4B" "#B63C37"
```

< ロ > < 同 > < 回 > < 回 > < 回 > <

### C: Summaries of spectra

e.g. irradiance (integral over wavelengths): total

$$I_{\lambda_{\min}...\lambda_{\max}} = \int_{\lambda=\lambda_{\min}}^{\lambda=\lambda_{\max}} I(\lambda) d\lambda$$
(3)

```
irrad(sun.spct)
      Total
##
##
   269.1249
   attr(,"time.unit")
##
##
   [1] "second"
  attr(,"radiation.unit")
##
##
   [1] "energy irradiance total"
```

< 口 > < 同 >

#### C: Summaries of spectra

e.g. irradiance (integral over wavelengths): ultraviolet-A

$$I_{\text{UV-A}} = \int_{\lambda=\text{start}(\text{UV-A})}^{\lambda=\text{end}(\text{UV-A})} I(\lambda) d\lambda$$
(4)

```
irrad(sun.spct, UVA())
##
    UVA.ISO
   27,98418
##
   attr(,"time.unit")
##
  [1] "second"
##
## attr(,"radiation.unit")
## [1] "energy irradiance total"
```

< 口 > < 同 >

#### — Examples

### C: Summaries of spectra

e.g. weighted irradiance (integral over wavelengths): CIE erythemal (human skin reddening)

$$I_{\text{CIE}} = \int_{\lambda=\text{start}(w_{\text{CIE}})}^{\lambda=\text{end}(w_{\text{CIE}})} I(\lambda) \times w_{\text{CIE}}(\lambda) d\lambda$$
(5)

```
irrad(sun.spct, CIE())
## CIE98.298.tr.lo
## 0.08181415
## attr(,"time.unit")
## [1] "second"
## attr(,"radiation.unit")
## [1] "energy irradiance total"
```

#### - Examples

### C: Summaries of spectra

e.g. irradiance for the band(s) defined according to wavelengths in nanometres

irrad(sun.spct, waveband(c(400, 700)))

## range.400.700
## 196.6343
## attr(,"time.unit")
## [1] "second"
## attr(,"radiation.unit")
## [1] "energy irradiance total"

irrad(sun.spct, split\_bands(c(400, 500, 600, 700)))

## range.400.500 range.500.600 range.600.700 ## 69.69042 68.48951 58.45435 ## attr(,"time.unit") ## [1] "second" ## attr(,"radiation.unit") ## [1] "energy irradiance total"

< ロ > < 同 > < 三 > < 三 > < 三 > <

#### - Examples

### C: Summaries of spectra

e.g. irradiance for the band(s) defined according to wavelengths in nanometres

irrad(sun.spct, waveband(c(400, 700)))

```
## range.400.700
## 196.6343
## attr(,"time.unit")
## [1] "second"
## attr(,"radiation.unit")
## [1] "energy irradiance total"
```

irrad(sun.spct, split\_bands(c(400, 500, 600, 700)))

```
## range.400.500 range.500.600 range.600.700
## 69.69042 68.48951 58.45435
## attr(,"time.unit")
## [1] "second"
## attr(,"radiation.unit")
## [1] "energy irradiance total"
```

< ロ > < 同 > < 回 > < 回 > < 回 >

#### C: Summaries of spectra

e.g. photon (quantum) irradiance for two bands and the corresponding photon ratio

```
q_irrad(sun.spct, list(Red(), Blue()))
```

Red.ISO Blue.ISO ## 0.0004511084 0.0001490288 ## attr(,"time.unit") ## [1] "second" ## ## attr(,"radiation.unit") ## [1] "photon irradiance total"

Image: A matrix

#### - Examples

#### C: Summaries of spectra

e.g. photon (quantum) irradiance for two bands and the corresponding photon ratio

```
q_irrad(sun.spct, list(Red(), Blue()))
```

```
## Red.ISO Blue.ISO
## 0.0004511084 0.0001490288
## attr(,"time.unit")
## [1] "second"
## attr(,"radiation.unit")
## [1] "photon irradiance total"
```

```
q_ratio(sun.spct, Red(), Blue())
## Red.ISO: Blue.ISO(q:q)
## 3.026988
## attr(,"radiation.unit")
## [1] "q:q ratio"
```

#### — Examples

### C: Summaries of spectra

e.g. simulating effect of a filter on irradiance

irrad(sun.spct, UVA())

```
## UVA.IS0
## 27.98418
## attr(,"time.unit")
## [1] "second"
## attr(,"radiation.unit")
## [1] "energy irradiance total"
```

irrad(sun.spct \* gg400.spct, UVA())

```
## UVA.IS0
## 1.100092
## attr(,"time.unit")
## [1] "second"
## attr(,"radiation.unit")
## [1] "energy irradiance total"
```

(日)

#### — Examples

### C: Summaries of spectra

e.g. simulating effect of a filter on irradiance

irrad(sun.spct, UVA())

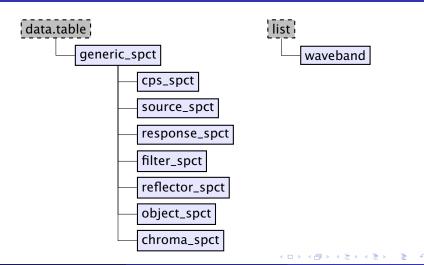
```
## UVA.IS0
## 27.98418
## attr(,"time.unit")
## [1] "second"
## attr(,"radiation.unit")
## [1] "energy irradiance total"
```

irrad(sun.spct \* gg400.spct, UVA())

```
## UVA.IS0
## 1.100092
## attr(,"time.unit")
## [1] "second"
## attr(,"radiation.unit")
## [1] "energy irradiance total"
```

< ロ > < 同 > < 回 > < 回 >

#### **Classes** For spectral data and for wavebands



© Pedro J. Aphalo (Univ. of Helsinki)

R4Photobiology

#### Classes Class of returned objects

#### A filtered light source is a light source, not a filter:

```
class(gg400.spct * sun.spct)
```

## [1] "source\_spct" "generic\_spct" "data.table" "data.frame"

#### An attenuated light source is still a light source:



A stack of two filters is still a filter:

```
class(gg400.spct * ug1.spct)
## [1] "filter_spct" "generic_spct" "data.table" "data.frame"
```

#### Attributes

getTimeUnit(sun.spct)

## [1] "second"

```
getTimeUnit(sun.daily.spct)
```

## [1] "day"

getTimeUnit(sun.daily.spct, force.duration = TRUE)

## [1] "86400s (~1 days)"

3

イロト 人間ト イヨト イヨト

#### Attributes

getTimeUnit(sun.spct)

## [1] "second"

getTimeUnit(sun.daily.spct)

## [1] "day"

getTimeUnit(sun.daily.spct, force.duration = TRUE)

```
## [1] "86400s (~1 days)"
```

3

イロト 人間ト イヨト イヨト

### Attributes

getTimeUnit(sun.spct)

## [1] "second"

getTimeUnit(sun.daily.spct)

## [1] "day"

getTimeUnit(sun.daily.spct, force.duration = TRUE)

```
## [1] "86400s (~1 days)"
```

イロト 不得 トイヨト イヨト 二日

### Comments in data objects

Automatic comment in object after acquisition + calibration using R

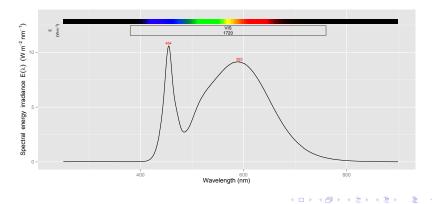
```
load("Lumitronix.cal.spcts.Rda")
cat(comment(Lumitronix_4cm.cal.spct))
## Spectrometer: MayaPro2000 s/n MAYP11278
## Bench with grating HC1, filter 000 and slit 010s
  Measured on 2015-05-19 18:23:34
##
## processed on 2015-05-19 with MayaCalc ver 3.2.3
## using HDR: TRUE, using NR: FALSE, method: full
   calibration dated (automatic): 2014-10-15
##
## number of scans
                     : 695.
                                              86
## integration times (ms) : 7.20, 58.53
## total times (s) : 5.0040, 5.0334
## max counts : 28410 out of 64000 (44%)
## LOW STGNAL!
```

#### Under the hood

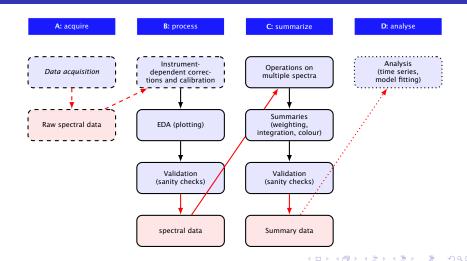
#### Plot same spectral data

Automatically generated axis labels and summaries: Units are 'known'

plot(Lumitronix\_4cm.cal.spct, w.band = VIS())



# Data flow



© Pedro J. Aphalo (Univ. of Helsinki)

R4Photobiology

UseR!2015, Aalborg, July 2015 30 / 33

# Where to find the suite of packages

- Web site at http://www.r4photobiology.info/
  - Installation instructions
  - Posts (also) RSS feed for package updates
  - Some other related posts and pages
  - Book/handbook early draft coming soon
- Packages not yet submitted to CRAN
- Own repository with CRAN-like structure at http://www.r4photobiology.info/R
- Git repository at Bitbucket http://bitbucket.org/aphalo/
- Other resources at http://www.uv4plants.org/

# Where to find the suite of packages

#### mailto:pedro.aphalo@helsinki.fi

Web site at http://www.r4photobiology.info/

- Installation instructions
- Posts (also) RSS feed for package updates
- Some other related posts and pages
- Book/handbook early draft coming soon
- Packages not yet submitted to CRAN
- Own repository with CRAN-like structure at http://www.r4photobiology.info/R
- Git repository at Bitbucket http://bitbucket.org/aphalo/
- Other resources at http://www.uv4plants.org/

# Where to find the suite of packages

#### mailto:pedro.aphalo@helsinki.fi

Web site at http://www.r4photobiology.info/

#### Installation instructions

- Posts (also) RSS feed for package updates
- Some other related posts and pages
- Book/handbook early draft coming soon
- Packages not yet submitted to CRAN
- Own repository with CRAN-like structure at http://www.r4photobiology.info/R
- Git repository at Bitbucket http://bitbucket.org/aphalo/
- Other resources at http://www.uv4plants.org/

### Where to find the suite of packages

- Web site at http://www.r4photobiology.info/
  - Installation instructions
  - Posts (also) RSS feed for package updates
  - Some other related posts and pages
  - Book/handbook early draft coming soon
- Packages not yet submitted to CRAN
- Own repository with CRAN-like structure at http://www.r4photobiology.info/R
- Git repository at Bitbucket http://bitbucket.org/aphalo/
- Other resources at http://www.uv4plants.org/

### Where to find the suite of packages

- Web site at http://www.r4photobiology.info/
  - Installation instructions
  - Posts (also) RSS feed for package updates
  - Some other related posts and pages
  - Book/handbook early draft coming soon
- Packages not yet submitted to CRAN
- Own repository with CRAN-like structure at http://www.r4photobiology.info/R
- Git repository at Bitbucket http://bitbucket.org/aphalo/
- Other resources at http://www.uv4plants.org/

# Where to find the suite of packages

- Web site at http://www.r4photobiology.info/
  - Installation instructions
  - Posts (also) RSS feed for package updates
  - Some other related posts and pages
  - Book/handbook early draft coming soon
- Packages not yet submitted to CRAN
- Own repository with CRAN-like structure at http://www.r4photobiology.info/R
- Git repository at Bitbucket http://bitbucket.org/aphalo/
- Other resources at http://www.uv4plants.org/

### Where to find the suite of packages

- Web site at http://www.r4photobiology.info/
  - Installation instructions
  - Posts (also) RSS feed for package updates
  - Some other related posts and pages
  - Book/handbook early draft coming soon
- Packages not yet submitted to CRAN
- Own repository with CRAN-like structure at http://www.r4photobiology.info/R
- Git repository at Bitbucket http://bitbucket.org/aphalo/
- Other resources at http://www.uv4plants.org/

### Where to find the suite of packages

- Web site at http://www.r4photobiology.info/
  - Installation instructions
  - Posts (also) RSS feed for package updates
  - Some other related posts and pages
  - Book/handbook early draft coming soon
- Packages not yet submitted to CRAN
- Own repository with CRAN-like structure at http://www.r4photobiology.info/R
- Git repository at Bitbucket http://bitbucket.org/aphalo/
- Other resources at http://www.uv4plants.org/

### Where to find the suite of packages

- Web site at http://www.r4photobiology.info/
  - Installation instructions
  - Posts (also) RSS feed for package updates
  - Some other related posts and pages
  - Book/handbook early draft coming soon
- Packages not yet submitted to CRAN
- Own repository with CRAN-like structure at http://www.r4photobiology.info/R
- Git repository at Bitbucket http://bitbucket.org/aphalo/
- Other resources at http://www.uv4plants.org/

# Where to find the suite of packages

- Web site at http://www.r4photobiology.info/
  - Installation instructions
  - Posts (also) RSS feed for package updates
  - Some other related posts and pages
  - Book/handbook early draft coming soon
- Packages not yet submitted to CRAN
- Own repository with CRAN-like structure at http://www.r4photobiology.info/R
- Git repository at Bitbucket http://bitbucket.org/aphalo/
- Other resources at http://www.uv4plants.org/

#### Acknowledgements

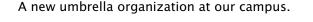




Members of my research group and collaborators (most of them listed below).

Members of the UV4Plants association: Lars Olof Björn, Andy McLeod, T. Matthew Robson, Titta Kotilainen, Luis O. Morales, Fang Wang, Marcel A. K. Jansen, Anders Lindfors and many others.







My employer.



For funding, decisions 252548, 16775.

EU for funding COST action FA0906 'UV4Growth' and our training events and meetings.

Equipment suppliers *Ocean Optics*, *sglux*, *Schott* for support and/or providing technical data on their products for inclusion as examples.

< ロ > < 同 > < 回 > < 回 > < 回 >

A very special thanks to the R and T<sub>E</sub>X communities for all the tools, packages, documentation and help that made developing this suite not only possible but also so enjoyable!

